

9. Supersymmetry

9.1 Introduction, concept

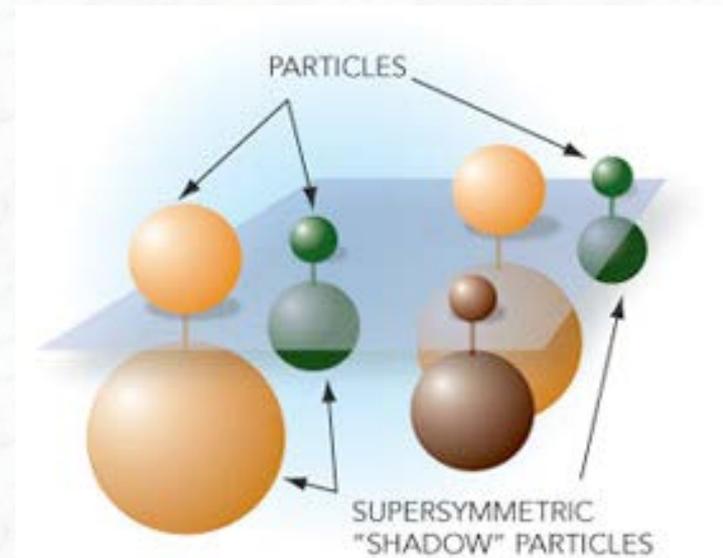
9.2 Motivation for SUSY

9.3 Breaking of Supersymmetry

9.4 Supersymmetry pre-LHC

9.5 Search for supersymmetry at the LHC

9.6 How can the parameters of a SUSY model be constrained



Why Physics Beyond the Standard Model?

1. Gravity is not yet incorporated in the Standard Model
2. Dark Matter not accommodated
3. Many open questions in the Standard Model
 - Hierarchy problem: m_W (100 GeV) \rightarrow m_{Planck} (10^{19} GeV)
 - Unification of couplings
 - Flavour / family problem
 -

All this calls for a **more fundamental theory** of which the Standard Model is a low energy approximation \rightarrow **New Physics**

Candidate theories: Supersymmetry
Extra Dimensions
.....

Many extensions predict new physics at the TeV scale !!

Strong motivation for LHC, mass reach \sim 3 TeV

9.1 Introduction: what is Supersymmetry (SUSY) ?

SUSY is an extension of the Standard Model (since ~ 1970) that introduces a **new symmetry between fermions and bosons**:

Spin 1/2 matter particles (fermions) \square **Spin 1** force particles (bosons)

SUSY transformation (operator Q):

$$Q |\text{Fermion}\rangle \sim |\text{Boson}\rangle$$

$$Q |\text{Boson}\rangle \sim |\text{Fermion}\rangle$$

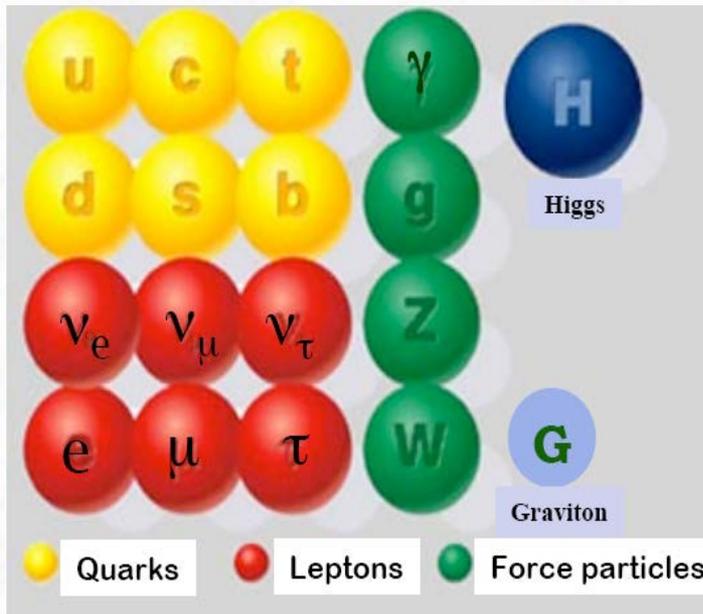
→ SUSY doubles the number of particles

SUSY is a symmetry, not a theory. A model (a Lagrangian) can possess supersymmetry as a property.

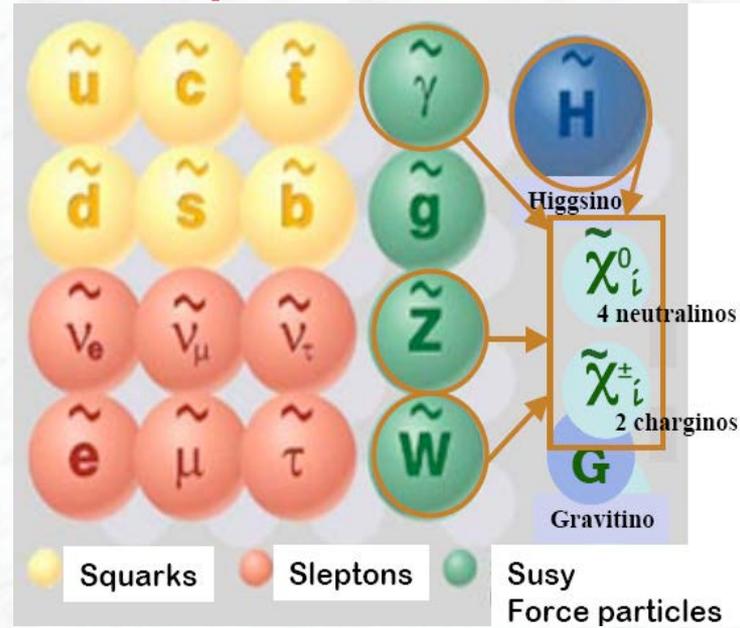
Supersymmetry

Extends the Standard Model by predicting a new symmetry
 Spin 1/2 matter particles (fermions) \Leftrightarrow Spin 1 force carriers (bosons)

Standard Model particles



SUSY particles



Standard Model and supersymmetry partners can be distinguished by a discrete Quantum number: R-parity:

$$R = (-1)^{3B+L+2S} = \begin{matrix} +1 & \text{SM particles} \\ -1 & \text{SUSY particles} \end{matrix}$$

The SUSY particle spectrum

	Particle	Sparticle (corresp. SUSY particle)
Spin-1/2	{ quarks (L&R) leptons (L&R) neutrinos (L)	

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Spin-1/2	<ul style="list-style-type: none"> quarks (L&R) leptons (L&R) neutrinos (L) 	<ul style="list-style-type: none"> squarks (L&R) sleptons (L&R) sneutrinos (L)

Spin-0

	Particle	Sparticle (corresp. SUSY particle)
Spin-1/2	<ul style="list-style-type: none"> quarks (L&R) leptons (L&R) neutrinos (L) 	<ul style="list-style-type: none"> squarks (L&R) sleptons (L&R) sneutrinos (L)
Spin-1	<ul style="list-style-type: none"> B W⁰ 	<ul style="list-style-type: none"> γ Z⁰ W[±] gluon

Spin-0

	Particle	Sparticle (corresp. SUSY particle)
Spin-1/2	<ul style="list-style-type: none"> quarks (L&R) leptons (L&R) neutrinos (L) 	<ul style="list-style-type: none"> squarks (L&R) sleptons (L&R) sneutrinos (L)
Spin-1	<ul style="list-style-type: none"> B W^0 <ul style="list-style-type: none"> γ Z^0 W^\pm gluon 	<ul style="list-style-type: none"> Bino Wino⁰ Wino[±] gluino

Spin-0

Spin-1/2

	Particle	Sparticle (corresp. SUSY particle)
Spin-1/2	{ quarks (L&R) leptons (L&R) neutrinos (L) }	{ squarks (L&R) sleptons (L&R) sneutrinos (L) } Spin-0
Spin-1	{ { B W ⁰ } { γ Z ⁰ W [±] gluon } }	{ Bino Wino ⁰ Wino [±] gluino } Spin-1/2
Spin-0	{ Higgs (H_1^1 H_1^2) (H_2^1 H_2^2) }	

Extended Higgs sector: **2 complex Higgs doublets**

→ Degrees of freedom: 8 - 3 (Goldstone bosons) = **5 Higgs bosons:** h^0, H^0, A^0, H^\pm

	Particle	Sparticle (corresp. SUSY particle)
Spin-1/2	quarks (L&R) leptons (L&R) neutrinos (L)	squarks (L&R) sleptons (L&R) sneutrinos (L)
Spin-1	$\left. \begin{array}{l} B \\ W^0 \end{array} \right\} \left\{ \begin{array}{l} \gamma \\ Z^0 \\ W^\pm \\ \text{gluon} \end{array} \right.$	Bino Wino ⁰ Wino [±] gluino
Spin-0	Higgs $\begin{pmatrix} H_1^1 \\ H_1^2 \end{pmatrix} \quad \begin{pmatrix} H_2^1 \\ H_2^2 \end{pmatrix}$	Higgsinos $\begin{pmatrix} \bar{H}_1^1 \\ \bar{H}_1^2 \end{pmatrix} \quad \begin{pmatrix} \tilde{H}_2^1 \\ \tilde{H}_2^2 \end{pmatrix}$

Spin-0

Spin-1/2

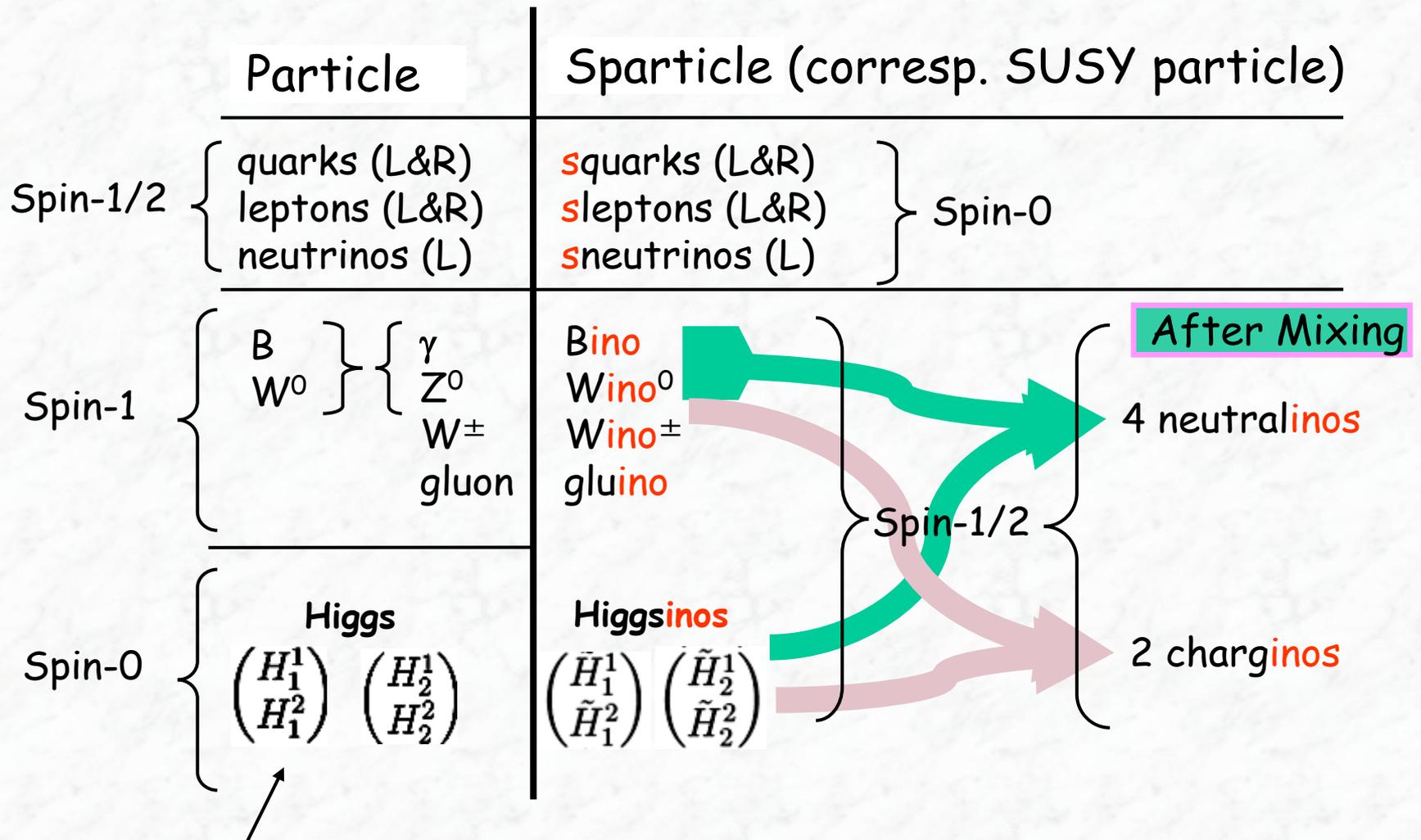
Extended Higgs sector: **2 complex Higgs doublets**
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Spin-1/2	quarks (L&R) leptons (L&R) neutrinos (L)	squarks (L&R) sleptons (L&R) sneutrinos (L)
Spin-1	B W^0	Bino $Wino^0$ $Wino^\pm$ gluino
Spin-0	Higgs $\begin{pmatrix} H_1^1 \\ H_1^2 \end{pmatrix}$ $\begin{pmatrix} H_2^1 \\ H_2^2 \end{pmatrix}$	Higgsinos $\begin{pmatrix} \tilde{H}_1^1 \\ \tilde{H}_1^2 \end{pmatrix}$ $\begin{pmatrix} \tilde{H}_2^1 \\ \tilde{H}_2^2 \end{pmatrix}$

The table is annotated with a large green arrow pointing from the Spin-1 and Spin-0 sparticle columns to the text "4 neutralinos" under the heading "After Mixing". A bracket labeled "Spin-1/2" encompasses the Bino, Winos, and gluino in the Spin-1 row, and the Higgsinos in the Spin-0 row.

Extended Higgs sector: 2 complex Higgs doublets

→ Degrees of freedom: 8 - 3 (Goldstone bosons) = 5 Higgs bosons: h^0, H^0, A^0, H^\pm



Extended Higgs sector: 2 complex Higgs doublets

→ Degrees of freedom: 8 - 3 (Goldstone bosons) = 5 Higgs bosons: h^0, H^0, A^0, H^\pm

- Physical **neutralinos** and **charginos** are **mixtures** of **Wino, Bino, Higgsinos**
- Charginos:

$$\begin{pmatrix} \chi_1^+ \\ \chi_2^+ \end{pmatrix} = \begin{pmatrix} M_2 & \sqrt{2}m_W \sin \beta \\ \sqrt{2}m_W \cos \beta & \mu \end{pmatrix} \begin{pmatrix} \tilde{W}^+ \\ \tilde{H}^+ \end{pmatrix}$$

- Neutralinos:

$$\begin{pmatrix} \chi_1^0 \\ \chi_2^0 \\ \chi_3^0 \\ \chi_4^0 \end{pmatrix} = \begin{pmatrix} M_1 & 0 & -m_Z c_\beta s_W & m_Z s_\beta s_W \\ 0 & M_2 & m_Z c_\beta c_W & -m_Z s_\beta c_W \\ -m_Z c_\beta s_W & m_Z c_\beta c_W & 0 & -\mu \\ m_Z s_\beta s_W & -m_Z s_\beta c_W & -\mu & 0 \end{pmatrix} \begin{pmatrix} \tilde{B} \\ \tilde{W}^3 \\ \tilde{H}_1^0 \\ \tilde{H}_2^0 \end{pmatrix}$$

Mass eigenstates depend on:

$M_1, M_2, \tan \beta, \mu$ SUSY masses and breaking parameters

$m_Z, \sin^2 \theta_W$ EWSB (mixing: $B^0, W^0 \rightarrow Z, \gamma$)

Supersymmetric particles carry the same quantum numbers (hypercharge, weak isospin, colour) as the Standard Model partners

That means they have the same interaction vertices

Names		spin 0	spin 1/2	$SU(3)_C, SU(2)_L, U(1)_Y$
squarks, quarks ($\times 3$ families)	Q	$(\tilde{u}_L \ \tilde{d}_L)$	$(u_L \ d_L)$	$(\mathbf{3}, \mathbf{2}, \frac{1}{6})$
	\bar{u}	\tilde{u}_R^*	u_R^\dagger	$(\bar{\mathbf{3}}, \mathbf{1}, -\frac{2}{3})$
	\bar{d}	\tilde{d}_R^*	d_R^\dagger	$(\bar{\mathbf{3}}, \mathbf{1}, \frac{1}{3})$
sleptons, leptons ($\times 3$ families)	L	$(\tilde{\nu} \ \tilde{e}_L)$	$(\nu \ e_L)$	$(\mathbf{1}, \mathbf{2}, -\frac{1}{2})$
	\bar{e}	\tilde{e}_R^*	e_R^\dagger	$(\mathbf{1}, \mathbf{1}, 1)$
Higgs, higgsinos	H_u	$(H_u^+ \ H_u^0)$	$(\tilde{H}_u^+ \ \tilde{H}_u^0)$	$(\mathbf{1}, \mathbf{2}, +\frac{1}{2})$
	H_d	$(H_d^0 \ H_d^-)$	$(\tilde{H}_d^0 \ \tilde{H}_d^-)$	$(\mathbf{1}, \mathbf{2}, -\frac{1}{2})$

Table 1.1: Chiral supermultiplets in the Minimal Supersymmetric Standard Model. The spin-0 fields are complex scalars, and the spin-1/2 fields are left-handed two-component Weyl fermions.

Names	spin 1/2	spin 1	$SU(3)_C, SU(2)_L, U(1)_Y$
gluino, gluon	\tilde{g}	g	(8, 1, 0)
winos, W bosons	$\tilde{W}^\pm \tilde{W}^0$	$W^\pm W^0$	(1, 3, 0)
bino, B boson	\tilde{B}^0	B^0	(1, 1, 0)

Gauge supermultiplets in the Minimal Supersymmetric Standard Model.

R parity

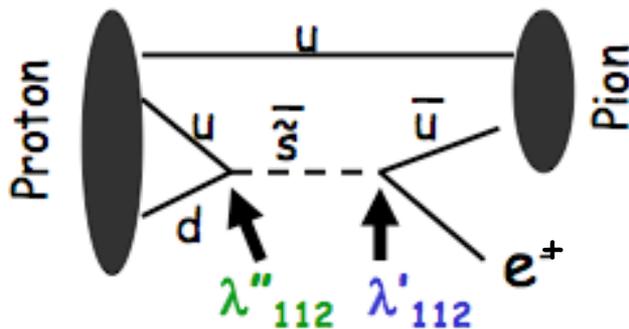
- New terms in Lagrangian:

$$W_{RPV} = \frac{1}{2} (\lambda LLE + \lambda' LQD + \lambda'' UDD) + \mu LH$$

↑
↑
↑
↑

L-violating
B-violating
L-violating

Problem: These couplings lead to **proton decay**



Unacceptably high rate compared to experimental limits
(proton lifetime $> 10^{33}$ years)

→ Strong limits on product of couplings

- Introduce multiplicative quantum number:

$$R_p = (-1)^{3(B-L)+2S}$$

= +1 for SM particles
 = -1 for SUSY particles

- Impose **R_p conservation**:
 - Sparticles produced in **pairs**
 - Lightest SUSY particle (**LSP**) **stable**

Experimental consequences of R-parity conservation:

- SUSY particles are **produced in pairs**
- **Lightest Supersymmetric Particle (LSP)** is stable.

LSP is only **weakly interacting**:

LSP $\equiv \chi^0_1$ (lightest neutralino, in many models)

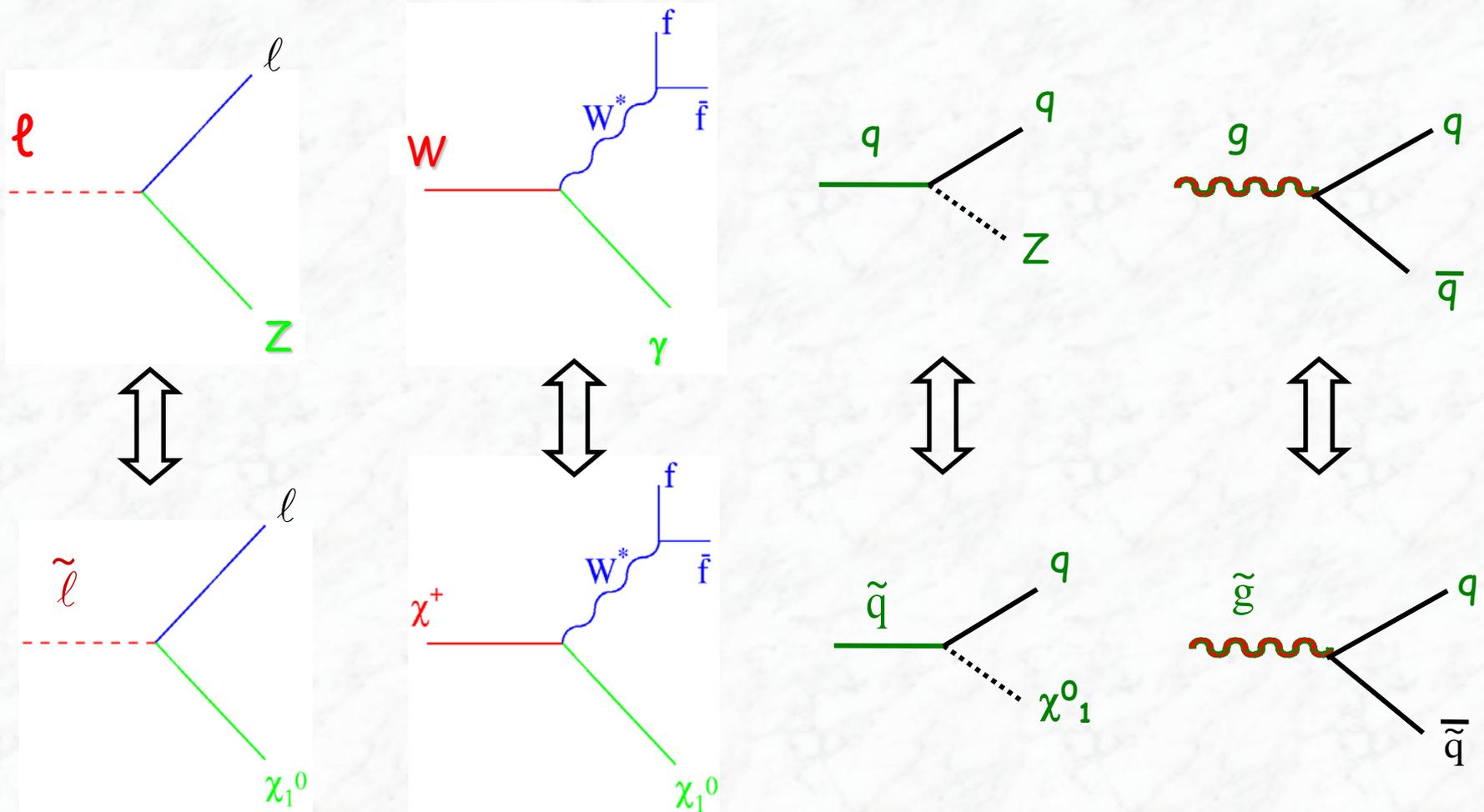
→ LSP behaves like a ν → it escapes detection

→ E_T^{miss} (typical SUSY signature)

SUSY interactions, some examples

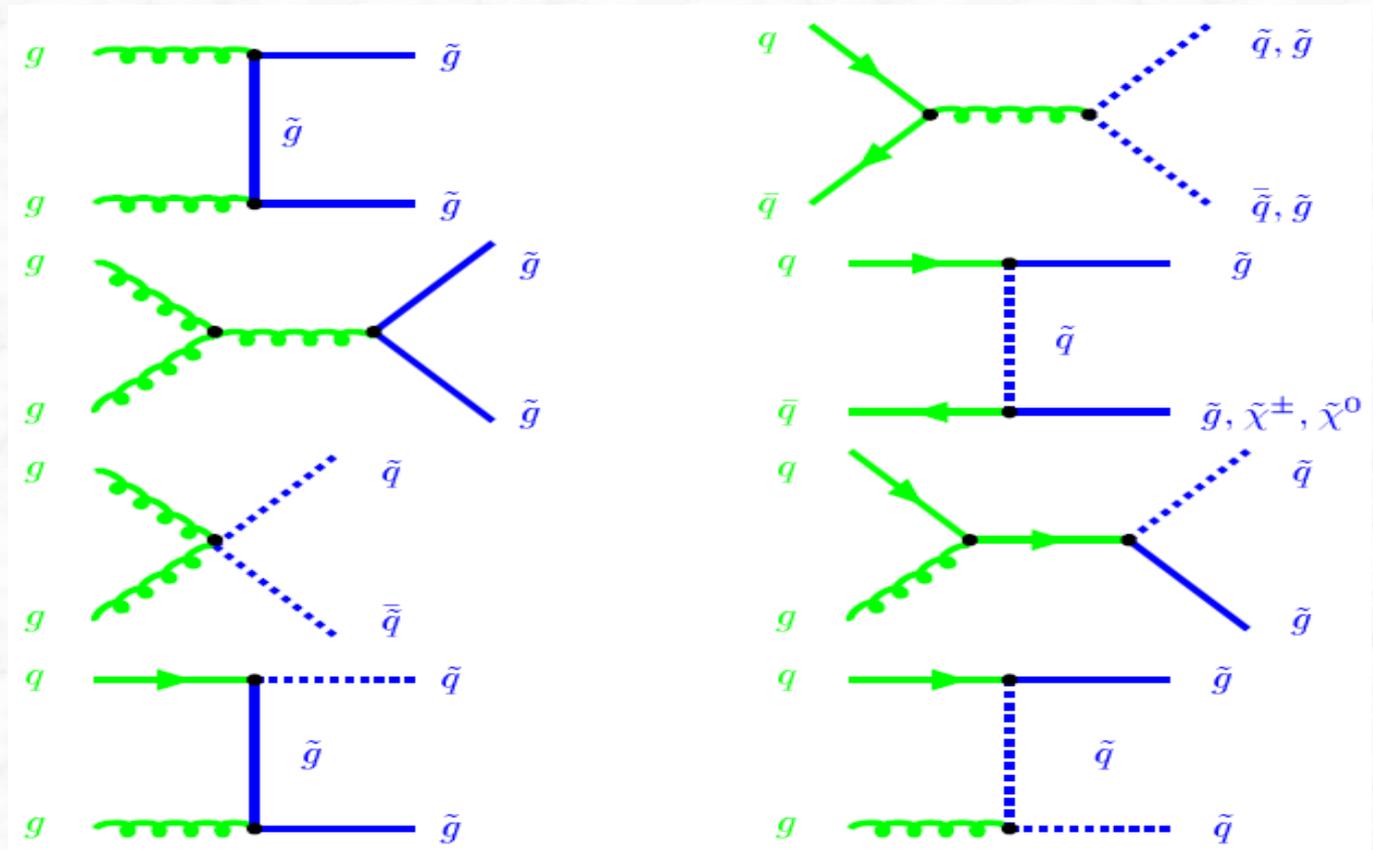
The coupling constants are the same as in SM (strong, electroweak)

“Recipe” : Obtain SUSY interactions by exchanging
at a vertex two SM legs by corresponding SUSY legs

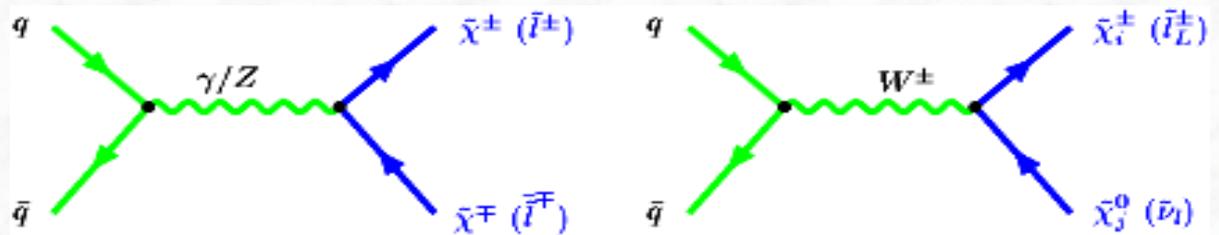


Sparticle production at the LHC

Quark-gluon fusion



Quark annihilation

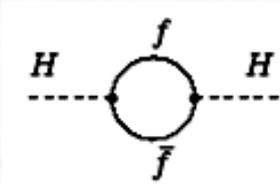


9.2 Motivation for supersymmetry

or what problems of the Standard Model does it solve ?

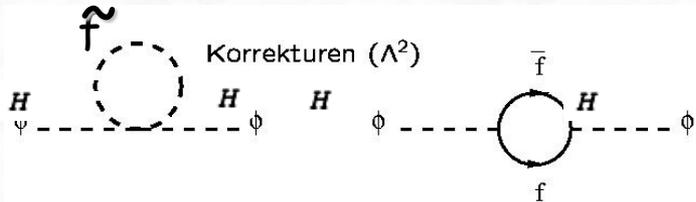
(i) The Hierarchy or naturalness problem

- In the Standard Model, quadratically divergent quantum corrections to the Higgs mass appear



$$\Delta m_H^2 = -\frac{|\lambda_f|^2}{8\pi^2} \Lambda_{UV}^2 + \dots$$

- The **symmetry between bosons and fermions**, which contribute with different sign (statistics), can cure this problem:

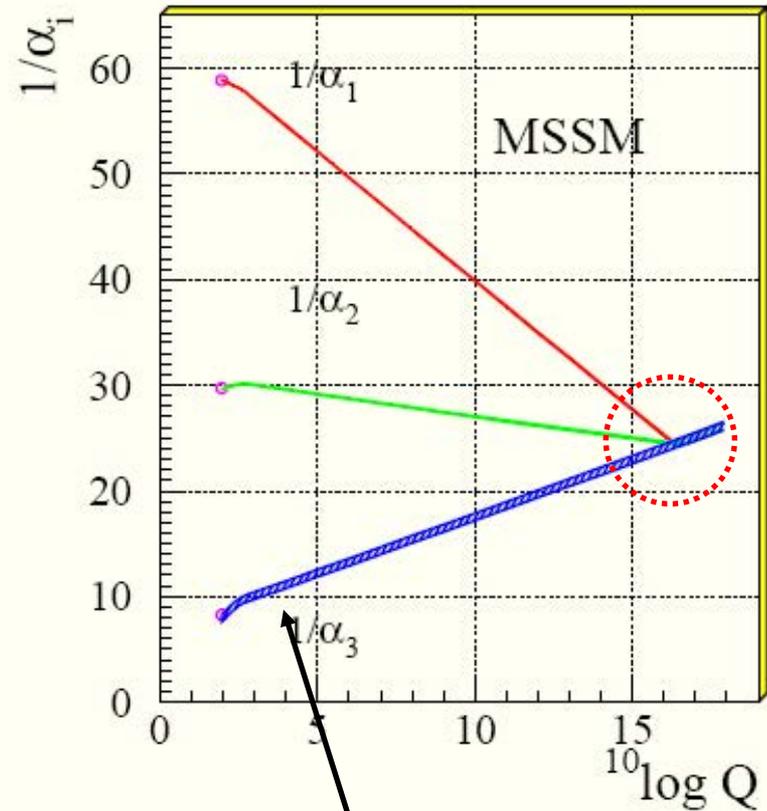
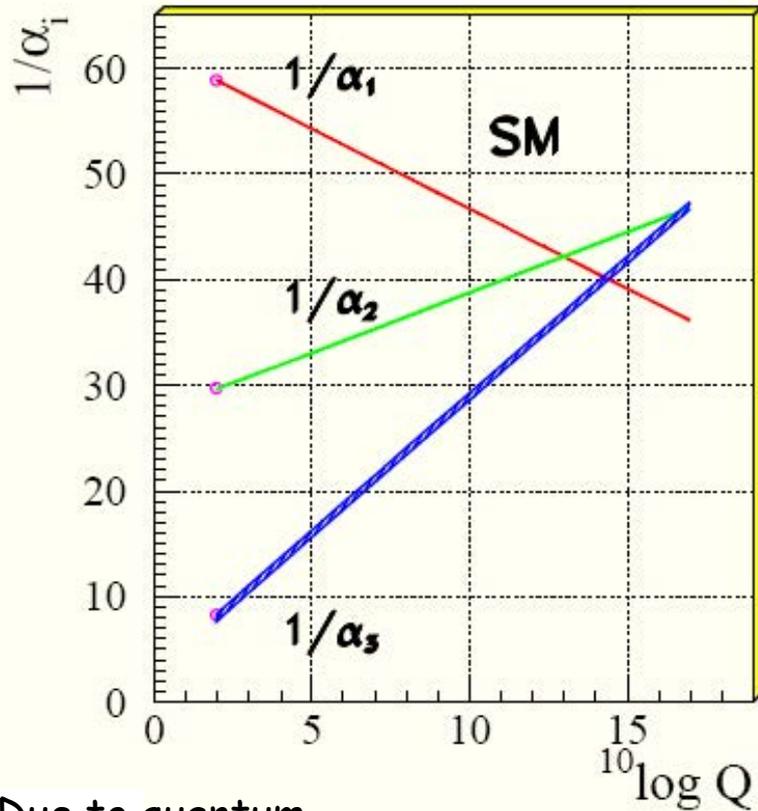


$$\Delta m_H^2 = \frac{\lambda_S}{16\pi^2} \left[\Lambda_{UV}^2 - 2m_S^2 \ln(\Lambda_{UV}/m_S) + \dots \right]$$

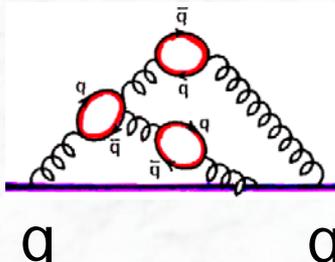
$$\Delta m_H = f(m_B^2 - m_f^2)$$

- terms cancel one-by-one if **SUSY were a perfect symmetry** (i.e. if $m(\text{particle}) = m(\text{sparticle})$). Since this is not the case, sparticles should not be too heavy ($m_{\text{SUSY}} < \sim 1 \text{ TeV}$).

(ii) The unification of the couplings of the three interactions seems possible in a SUSY model

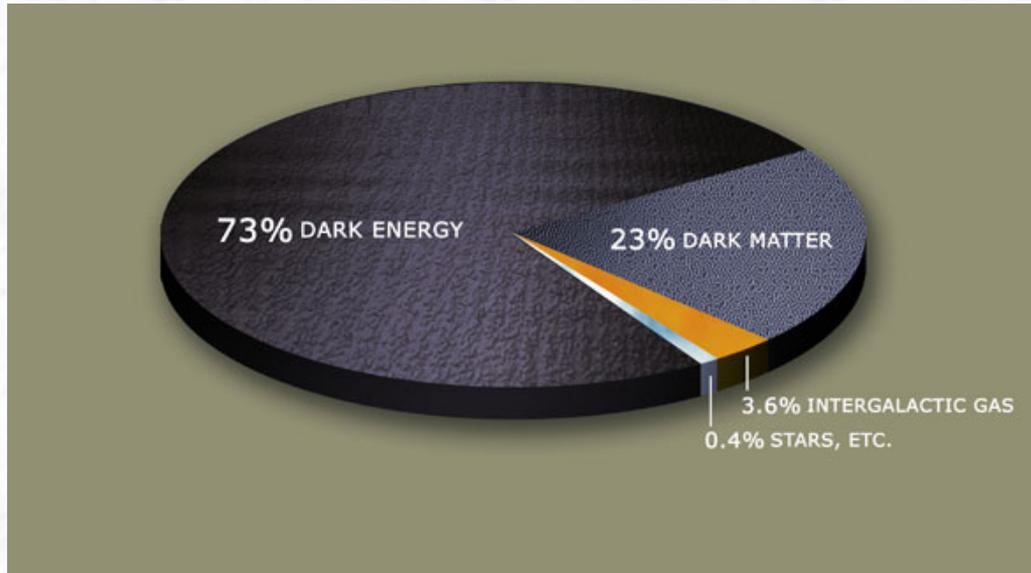


Due to quantum corrections, e.g.



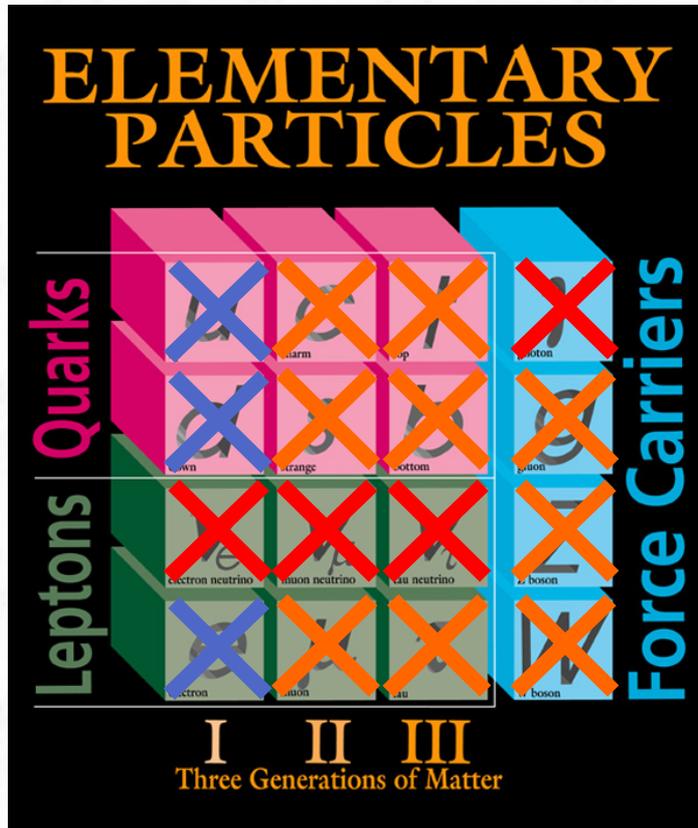
slope is changed due to contributions from SUSY particles

(ii) SUSY provides a candidate for the Dark Matter in the universe



Evidence from:

- Rotational curves of galaxies
- Gravitational lensing
- Cosmic microwave background (CMB)



Dark-Matter properties:

Gravitationally interacting

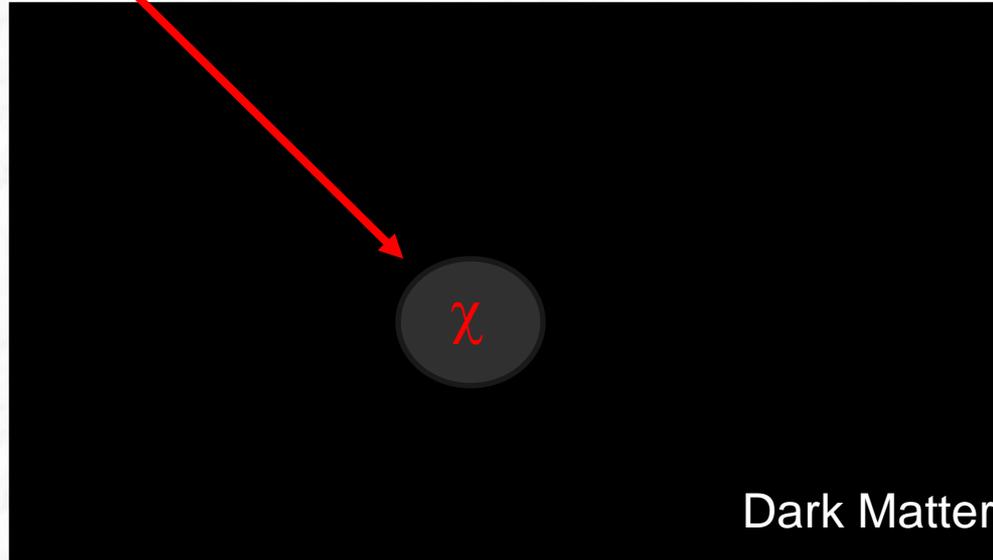
- Not short-lived
- Not hot
- Not baryonic, no el. magnetic interaction

→ Unambiguous evidence for new physics !

- SUSY has a **weakly interacting massive particle (WIMP)**, if R-parity is conserved:

the lightest supersymmetric particle

- **LSP** = lightest neutralino, gravitino (depending on SUSY model)



9.3 Breaking of Supersymmetry

- or the dark side of SUSY -

- Supersymmetry cannot be an exact symmetry, since we have not seen SUSY particles with masses equal to the Standard Model particles (e.g. no scalar electrons with masses of 511 keV)
- It is expected that supersymmetry is broken spontaneously, i.e. the underlying model should have a Lagrangian density that is invariant under supersymmetry, but a vacuum state that is not.
- The mechanism of SUSY breaking is not known
- Our ignorance is parametrized by introducing extra terms that break SUSY explicitly (so called soft SUSY breaking terms*) into the Lagrangian and consider it as an “effective” Lagrangian
- → about 105 parameters are introduced:
masses, couplings, mixing angles.....

*) only such terms are allowed that guarantee that no new quadratically divergent radiative corrections appear

A general parametrization comprises:

- Scalar mass terms:

$$m_{0_i}^2 \phi_i^2$$

- Gaugino mass terms

$$\frac{1}{2} M_a V_a^T C V_a$$

- Coupling terms of scalar particles

$$A_\lambda \lambda \phi^3 \quad , \quad B_\mu \mu \phi^2$$

It is often assumed that these supersymmetry breaking terms originate at some high scale (**Grand Unification scale (GUT scale)**, or gravity scale) maybe linked to some supergravity or superstring mechanism

It is furthermore often assumed that the soft supersymmetry-breaking terms are universal at the GUT or supergravity scale

The Minimal Supersymmetric Extension to the SM (MSSM)

It includes 105 additional parameters on top of the SM ones:

- SUSY violating mass terms for **each partner of the SM fermions**
- SUSY violating mass terms for **each partner of the SM bosons**
- Additional parameters on **the Higgs sector (5 Higgs bosons in total)**
- **Trilinear couplings** of all scalar particles

These parameters cannot be all independent:

- Or we would **have large FCNC, CP violation, etc.**
- Consistency with existing measurements indicate **some ordering principle is at play**

Bottom up approach - The phenomenological MSSM (pMSSM)

- Reasonable assumptions (motivated by current experimental results):
 - The mass matrices are diagonal and proportional to the identity matrix (no FCNC)
 - The trilinear couplings are proportional to the Yukawa couplings
 - No additional complex phase in the matrices (no additional CP violation)
 - Unified masses of first and second sfermions generations (from K^0 -anti- K^0 mixing)

Then one ends up with the so-called phenomenological MSSM, which has 19 parameters in addition to the SM parameters

$\tan \beta$: the ratio of the vev of the two-Higgs doublet fields.

M_A : the mass of the pseudoscalar Higgs boson

μ : the Higgs-higgsino mass parameter

M_1, M_2, M_3 : the bino, wino and gluino mass parameters.

$m_{\tilde{q}}, m_{\tilde{q}_R}, m_{\tilde{d}_R}, m_{\tilde{l}}, m_{\tilde{l}_R}$: first/second generation sfermion masses

$m_{\tilde{Q}}, m_{\tilde{l}_R}, m_{\tilde{b}_R}, m_{\tilde{\nu}_\tau}, m_{\tilde{\tau}_R}$: third generation sfermion masses

A_t, A_b, A_τ : third generation trilinear couplings.

Top-down approaches: Constrained MSSM models

It is often assumed that these supersymmetry breaking terms originate at some high scale (**Grand Unification scale (GUT scale)**, or gravity scale) maybe linked to some supergravity or superstring mechanism

It is furthermore often assumed that the soft supersymmetry-breaking terms are universal at the GUT or supergravity scale

The constrained MSSM models make strong assumptions about unification of SUSY parameters. They have been very popular for many years

But they really suffer from the LHC constraints..... With time, they are becoming less and less popular

If one assumes universality for all masses and couplings at the GUT scale, the following five parameters suffice to characterize the MSSM (Minimal Supersymmetric Standard Model)

Five parameters:

m_0 = common sfermion mass at the GUT scale

$m_{1/2} = M_1 = M_2$ (common gaugino mass at the GUT scale)

A_0 common coupling term

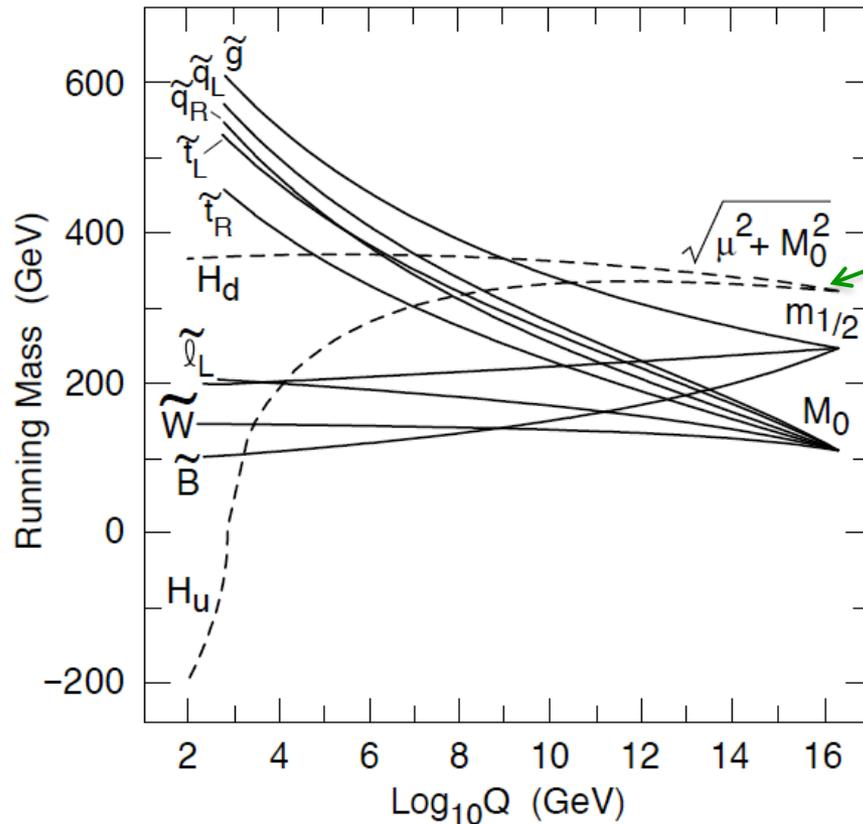
$\tan \beta$ ratio of vacuum expectation value of the two Higgs doublets

μ Higgs mass term

The parameters μ and $\tan \beta$ are related to the Higgs sector of the MSSM and determine this sector –together with one Higgs boson mass, e.g. m_A - at tree level completely.

This model is referred to as “constrained MSSM”

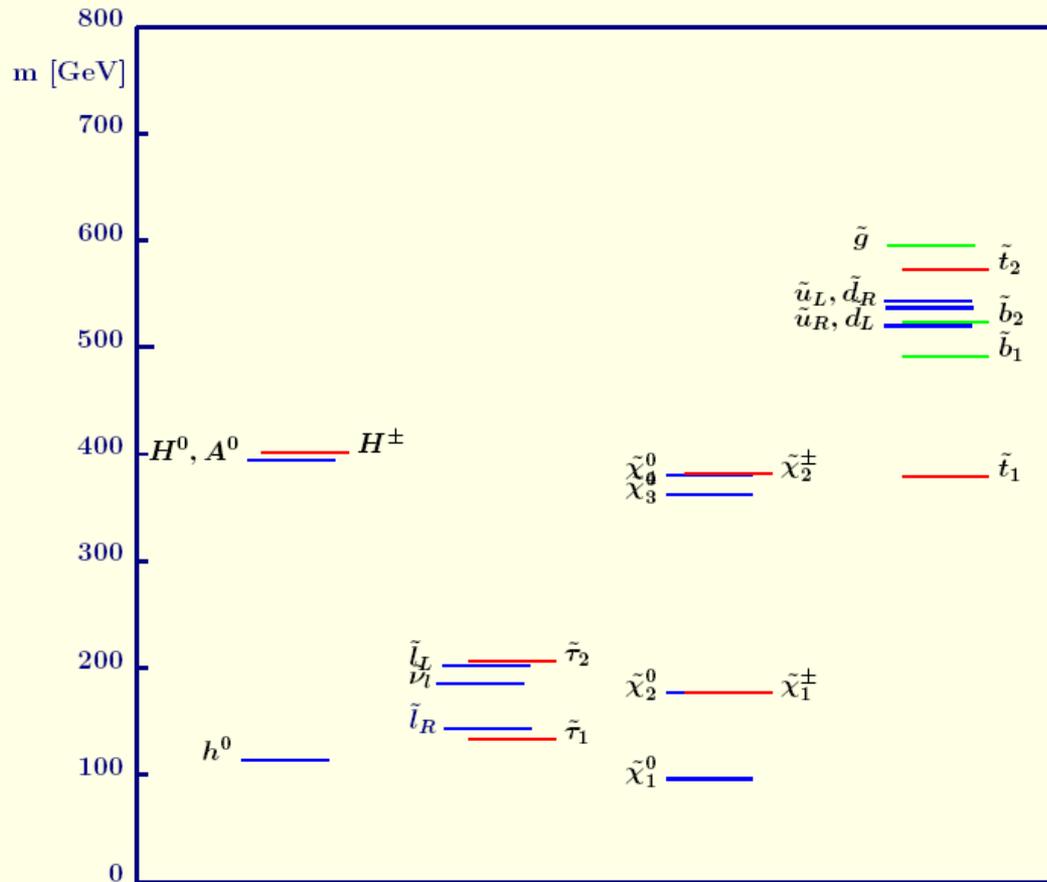
The energy dependence of the SUSY particle masses can be calculated (renormalization group equations)



In this example separate Higgs boson mass terms are used at the GUT scale

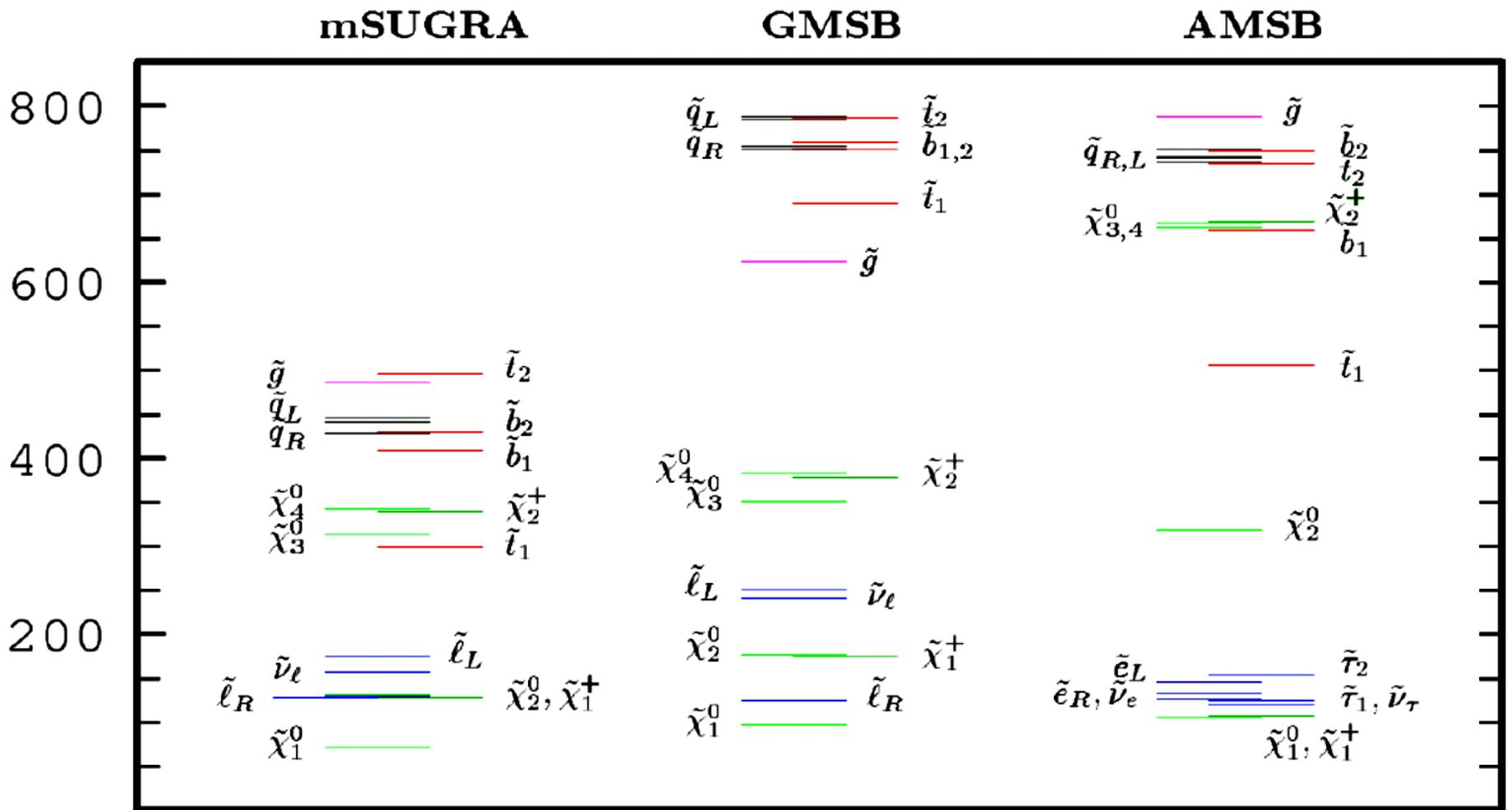
- In such models, squarks and/or gluinos are the heaviest sparticles
- Stops might be the lightest squarks
- Mixing effects (see later) can lead to mass splitting between t_L and t_R
- Sleptons / charginos / neutralinos are lightest sparticles

Example for a calculated mass spectrum at the electroweak scale:



- In many models, the lightest neutralino is the lightest SUSY particle (LSP)
- Lightest Higgs boson (h^0) might also be light, split from heavier Higgs particles
- Production of SUSY particles at the LHC is dominated by coloured squarks and gluinos

SUSY mass spectra for different SUSY breaking models:



mSUGRA: minimal SuperGRAvity model (gravity responsible for SUSY breaking)

GMSB: Gauge mediated SUSY Breaking (breaking via gauge interactions)

AMSB: Anomaly mediated SUSY Breaking

Sparticle Masses and Mixing

(i) Sfermions:

- Each flavour of charged leptons or quarks has both left- and right handed components, $f_{L,R}$ and these have separate spin-0 boson superpartners $f'_{L,R}$.
- The superpartners can mix, and the mass matrix in MSSM is given by:

$$M_{\tilde{f}}^2 \equiv \begin{pmatrix} m_{\tilde{f}LL}^2 & m_{\tilde{f}LR}^2 \\ m_{\tilde{f}LR}^2 & m_{\tilde{f}RR}^2 \end{pmatrix}$$

where the off-diagonal mixing terms take the general form (with m_f being the corresponding fermion mass).

$$m_{\tilde{f}L,R}^2 = m_f \left(A_f + \mu \frac{\tan \beta}{\cot \beta} \right) \quad \text{for } f = \begin{matrix} e, \mu, \tau, d, s, b \\ u, c, t \end{matrix}$$

- Since mixing effects are proportional to the SM fermion masses, they are in particular important for third generation sfermions. They are as well enhanced for large $\tan \beta$.
- Diagonalization of this mass matrix leads to the physical SUSY particle masses. Mixing might split the two states in mass, usually referred to as the lighter f'_1 and the heavier state f'_2 .

Sparticle Masses and Mixing

(ii) Charginos:

- Charginos are the supersymmetric partners of the W^\pm and charged Higgs bosons H^\pm .

Their masses are determined via the parameters: M_2 , μ and $\tan \beta$ appearing in the mass mixing matrix M_C :

$$-\frac{1}{2} (\tilde{W}^-, \tilde{H}^-) M_C \begin{pmatrix} \tilde{W}^+ \\ \tilde{H}^+ \end{pmatrix} + \text{herm.conj.}$$

$$M_C \equiv \begin{pmatrix} M_2 & \sqrt{2}m_W \sin \beta \\ \sqrt{2}m_W \cos \beta & \mu \end{pmatrix}$$

where M_2 is the unmixed SU(2) gaugino mass and μ is the Higgs mass parameter.

Sparticle Masses and Mixing

(ii) Neutralinos:

- Masses of neutralinos, the supersymmetric partners of the neutral fields

$$(\tilde{W}^3, \tilde{B}, \tilde{H}_2^0, \tilde{H}_1^0)$$

are given by the following mixing matrix:

$$m_N = \begin{pmatrix} M_2 & 0 & \frac{-g_2 v_2}{\sqrt{2}} & \frac{g_2 v_1}{\sqrt{2}} \\ 0 & M_1 & \frac{g' v_2}{\sqrt{2}} & \frac{-g' v_1}{\sqrt{2}} \\ \frac{-g_2 v_2}{\sqrt{2}} & \frac{g' v_2}{\sqrt{2}} & 0 & \mu \\ \frac{g_2 v_1}{\sqrt{2}} & \frac{-g' v_1}{\sqrt{2}} & \mu & 0 \end{pmatrix}$$

where M_1, M_2 is the unmixed U(1), SU(2) gaugino mass, μ is the Higgs mass parameter, and g_2 and g' are the gauge couplings of the SU(2) / U(1) group.

Higgs boson and stop quark

Because of the large top Yukawa coupling, the Higgs boson and the stop sector turn out to be heavily connected

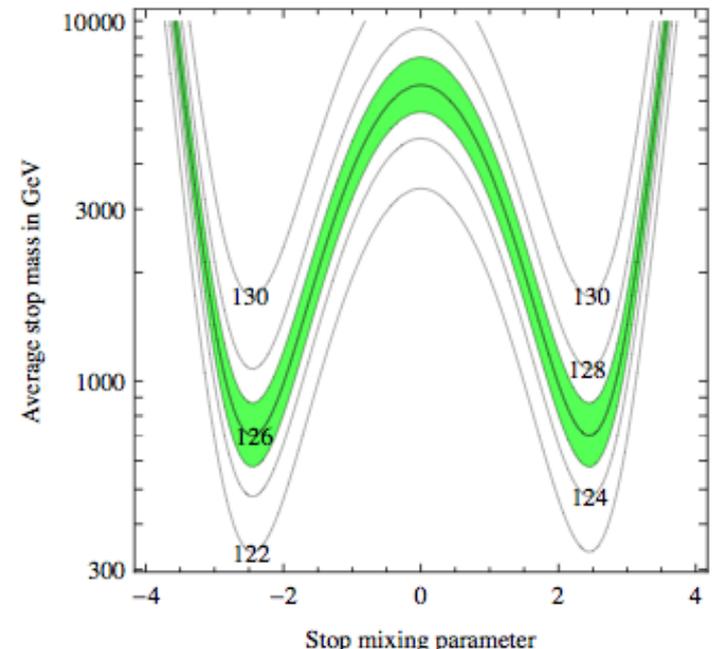
$$m_h^2 = m_Z^2 \cos^2 \beta + \frac{3y_t^2 m_t^2}{(4\pi)^2} \left[\log \left(\frac{m_S^2}{m_t^2} \right) + X_t^2 \left(1 - \frac{X_t^2}{12} \right) \right] + \dots$$

The natural Higgs mass is m_Z . However loop corrections are large

$$X_t = \left(A_t - \frac{\mu}{\tan \beta} \right)$$

$$M_S = \sqrt{m_{\tilde{t}_1} m_{\tilde{t}_2}}$$

$$M_{\tilde{t}}^2 = \begin{bmatrix} m_Q^2 + m_t^2 & m_t(A_t - \mu^*/\tan \beta) \\ m_t(A_t^* - \mu/\tan \beta) & m_U^2 + m_t^2 \end{bmatrix}$$



Link to the Dark Matter in the Universe ?

Parameters of the SUSY model \Rightarrow predictions for the relic density of dark matter

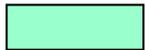
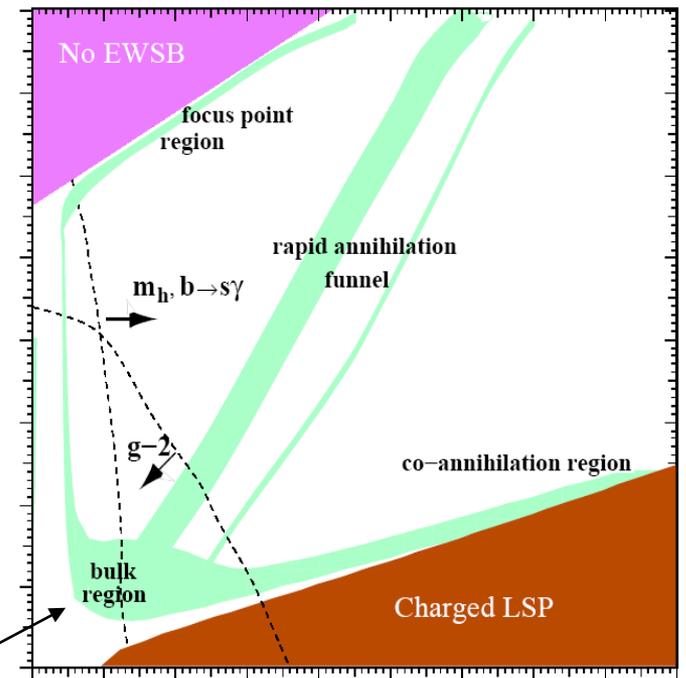
Interpretation in a simplified model

$$\rho_\chi \sim m_\chi n_\chi, \quad n_\chi \sim \frac{1}{\sigma_{ann}(\chi\chi \rightarrow \dots)}$$

cMSSM
(constrained Minimal Supersymmetric Standard Model)

Five parameters:

- $m_0, m_{1/2}$ particle masses at the GUT scale
- A_0 common coupling term
- $\tan \beta$ ratio of vacuum expectation value of the two Higgs doublets
- μ (sign μ) Higgs mass term



regions of parameter space which are consistent with the measured relic density of dark matter (WMAP,.....)

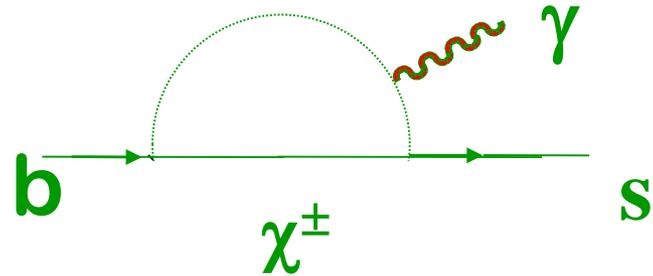
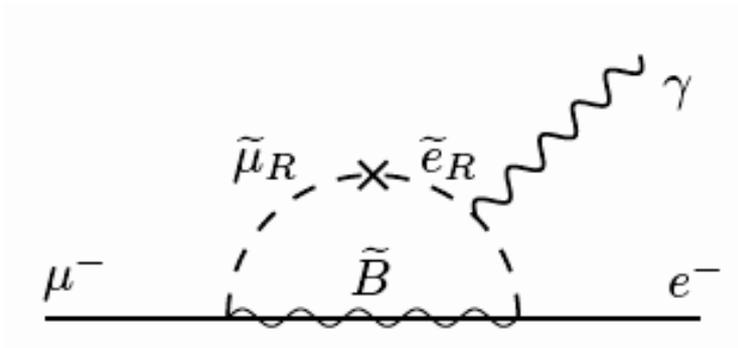
9.4 Supersymmetry pre-LHC

9.4.1 SUSY searches at LEP

9.4.2 SUSY searches at the Tevatron

Indirect SUSY searches

- Measure branching fractions of **rare decays** or search for **forbidden decays**.
→ Potentially enhanced by SUSY particles “in loops”:



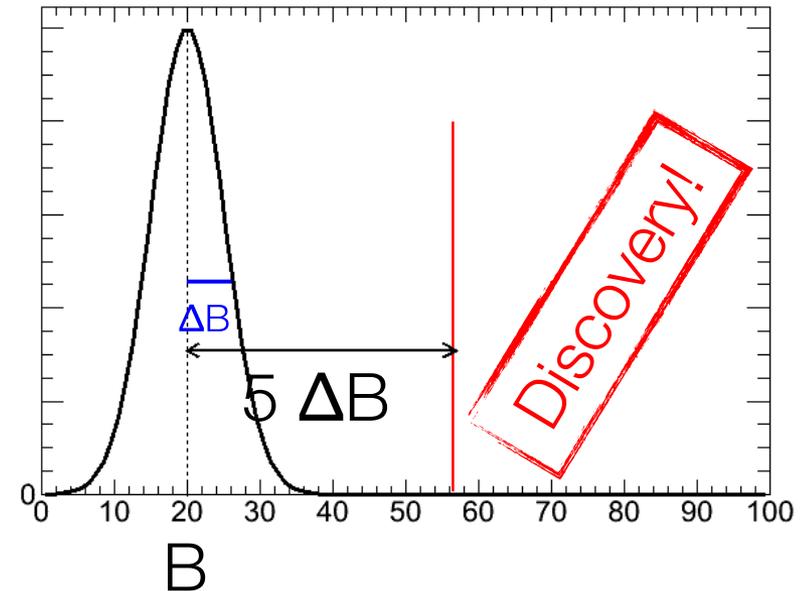
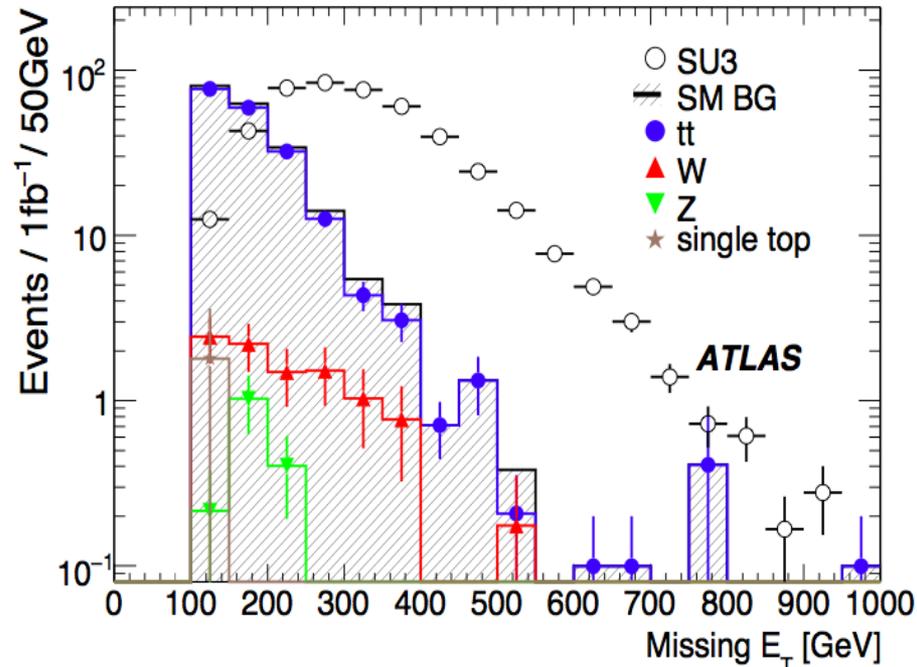
- Measurement of a_μ **anomalous magnetic moment** ($g_\mu - 2$): Brookhaven

Direct searches – general approach

- R-parity conserving Supersymmetry means **missing transverse momentum** in the final state

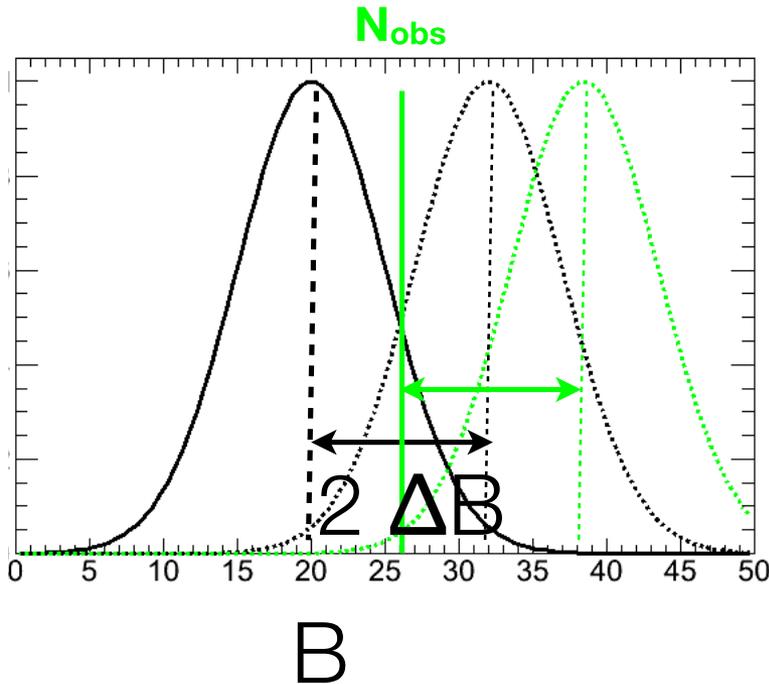
The simplest possible search strategy:
- A harsh cut on the missing transverse momentum will strongly reduce the number of expected events from known Standard Model Processes

→ Search for excess in high E_T^{miss} region



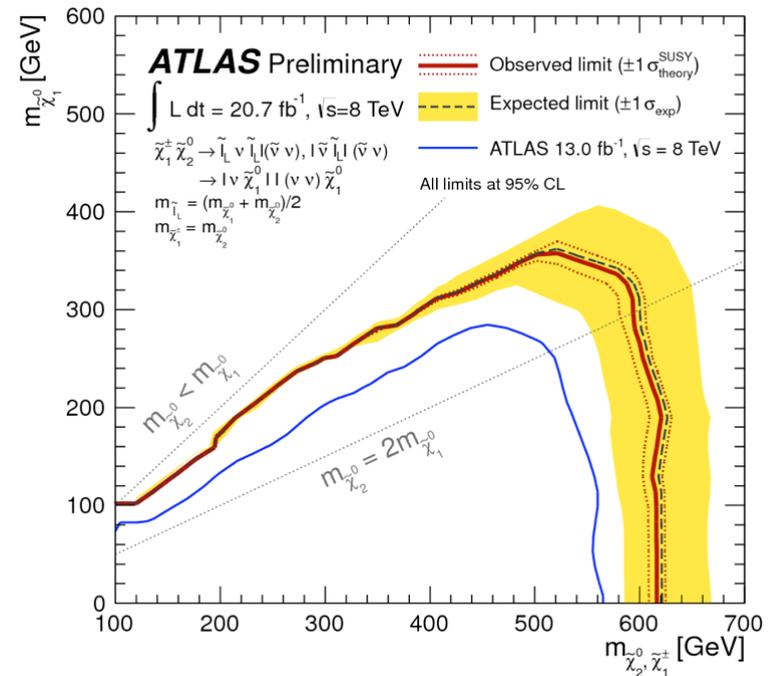
Direct searches – limit settings

- What if instead the number of observed events is consistent with the SM prediction?



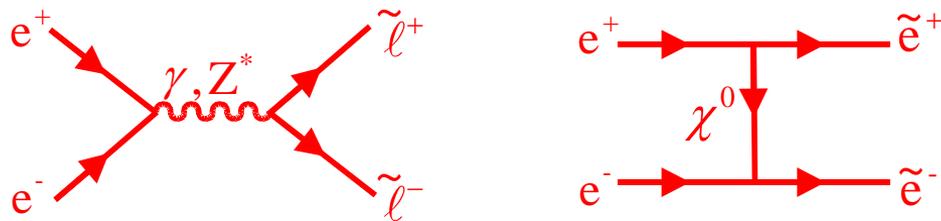
- For a given signal hypothesis, one can compute if the number of expected (observed) SM events differs from the expected (observed) SM + SUSY yield for more than 2 Gaussian equivalent sigmas

If yes, the model is excluded (either expected to be excluded or excluded)



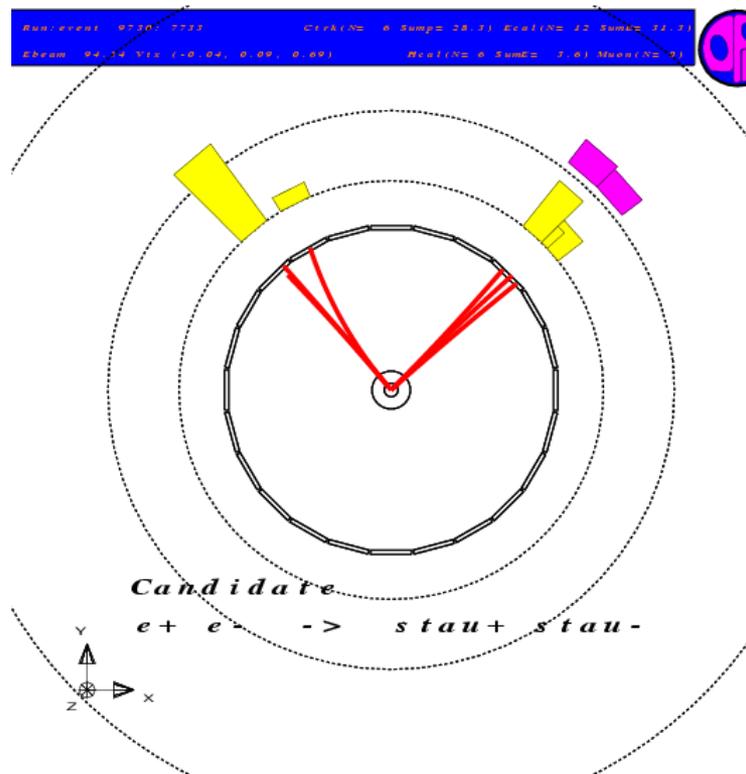
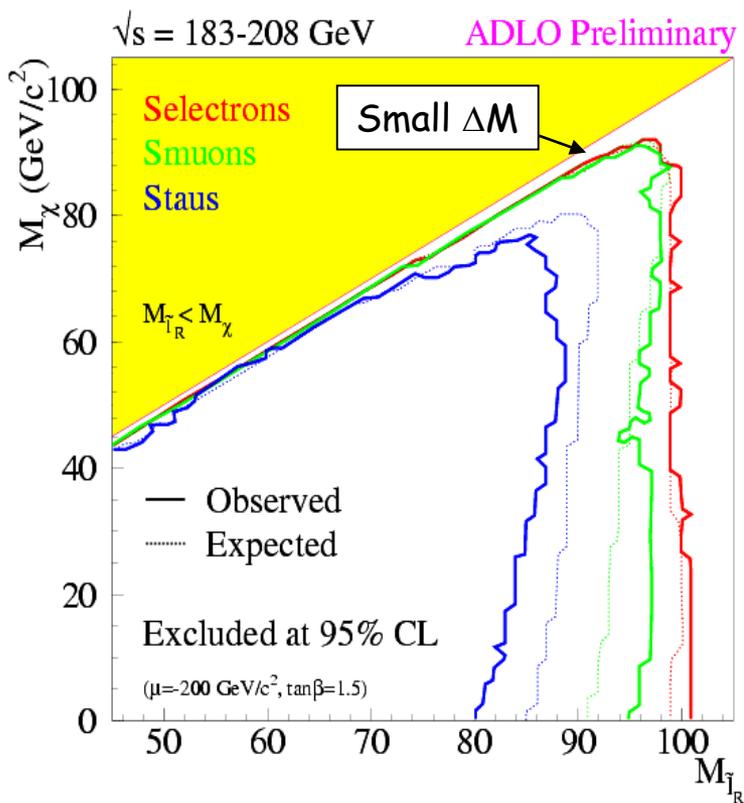
Direct searches for sleptons at LEP

- Clear domain of LEP are **Slepton & Chargino/Neutralino** searches (e^+e^- , $E_{cm} \sim 200$ GeV)



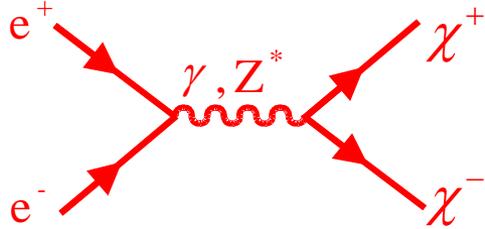
$$\tilde{l} \rightarrow l \chi^0_1 \quad (\text{2 leptons + missing E})$$

OPAL stau event candidate



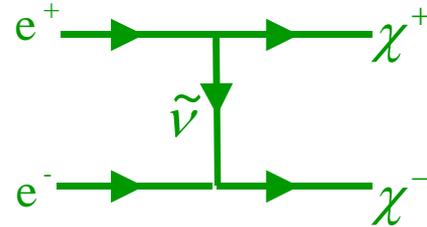
Direct searches for charginos at LEP

Large m_0 (\tilde{l} are heavy)



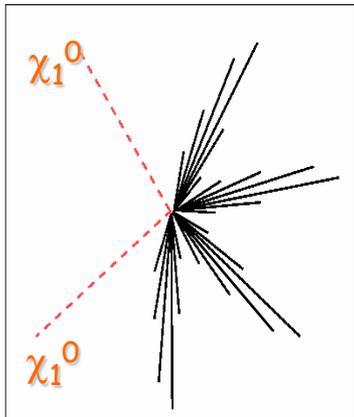
$$\chi^+\chi^- \rightarrow W^*\chi_1^0 W^*\chi_1^0$$

Small m_0 (\tilde{l} are light)

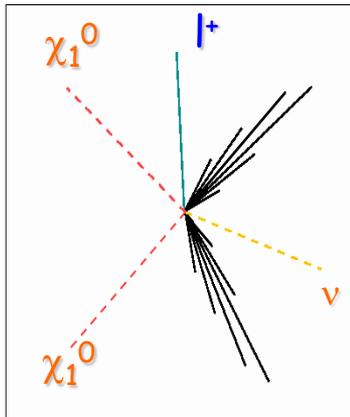


$$\chi^+\chi^- \rightarrow l^+\tilde{\nu} l^-\tilde{\nu} \rightarrow l^+\nu\chi_1^0 l^-\nu\chi_1^0$$

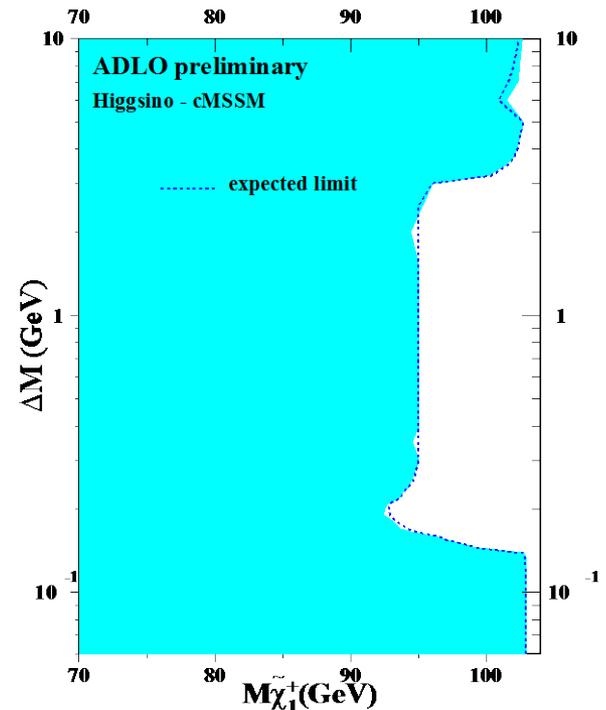
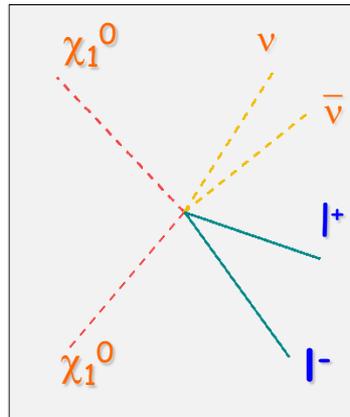
WW \square qqqq



WW \square l^+l^- qq



WW \square l^+l^- $\nu\bar{\nu}$



Main SM backgrounds

(WW, ZZ production) suppressed by requiring large missing mass or missing energy in the event.

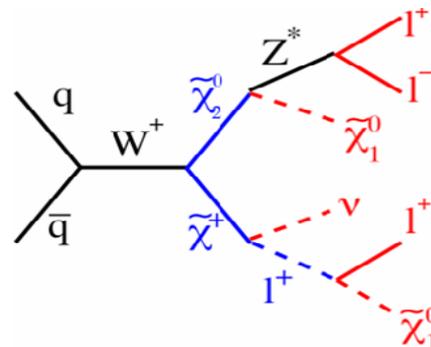
The two classical SUSY signatures at the Tevatron

1. Search for Squarks and Gluinos: **Jet + E_T^{miss}** signature
 produced via QCD processes



2. Search for Charginos and Neutralinos: **Multilepton + E_T^{miss}** signature
 produced via electroweak processes (associated production – not discussed here)

$$\tilde{\chi}_2^0 \tilde{\chi}_1^\pm \rightarrow l^\pm l^\mp l^\pm \tilde{\chi}_1^0 \tilde{\chi}_1^0 X$$



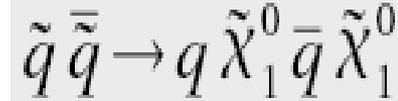


Search for Squarks and Gluinos

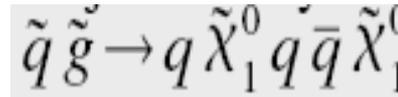


- Three different analyses, depending on squark / gluinos mass relations:

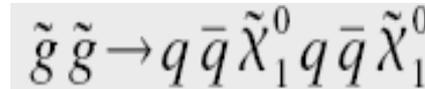
(i) dijet analysis
small m_0 , $m(\text{squark}) < m(\text{gluino})$



(ii) 3-jet analysis
intermediate m_0 $m(\text{squark}) \approx m(\text{gluino})$

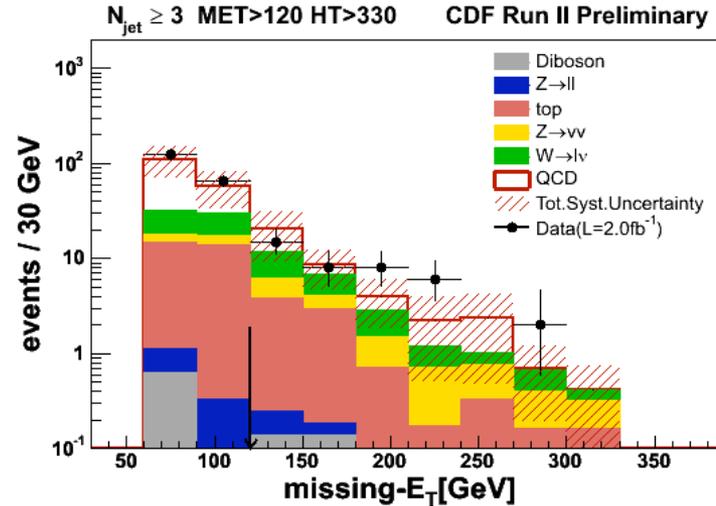
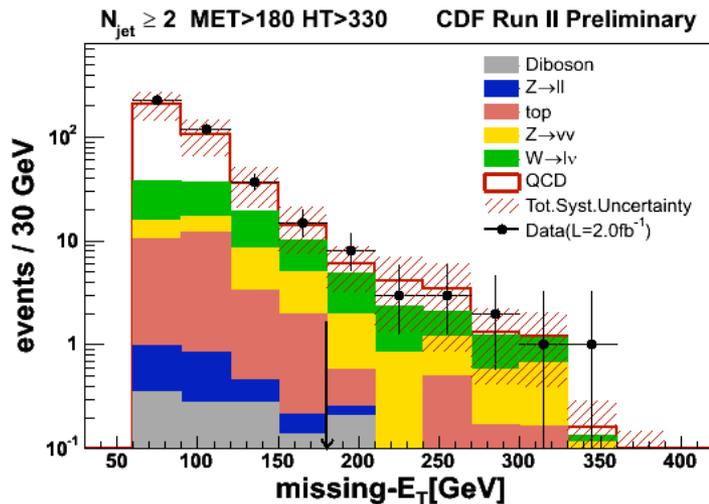


(iii) Gluino analysis
large m_0 , $m(\text{squark}) > m(\text{gluino})$



- **Main backgrounds:** $Z \rightarrow \nu\nu + \text{jets}$, $t\bar{t}$, $W + \text{jet production}$
- **Event selection:**
 - * require at least 2, 3 or 4 jets with $P_T > 60 / 40 / 30 / 20$ GeV
 - * veto on isolated electrons and muons
 - * isolation of E_T^{miss} and all jets
 - * optimization of the final cuts \rightarrow discriminating variables

Search for Squarks and Gluinos (cont.)



Expected background:

samples	2-jets	3-jets	4-jets
QCD	4.37 ± 2.01	13.34 ± 4.67	15.26 ± 7.60
top	1.35 ± 1.22	7.56 ± 3.85	22.14 ± 7.29
$Z \rightarrow \nu\nu + \text{jets}$	3.95 ± 1.09	5.39 ± 1.74	2.74 ± 0.95
$Z \rightarrow ll + \text{jets}$	0.09 ± 0.04	0.16 ± 0.11	0.14 ± 0.08
$W \rightarrow lv + \text{jets}$	6.08 ± 2.15	10.69 ± 3.84	7.68 ± 2.85
WW/WZ/ZZ	0.21 ± 0.19	0.35 ± 0.17	0.49 ± 0.34
tot SM	16 ± 5	37 ± 12	48 ± 17

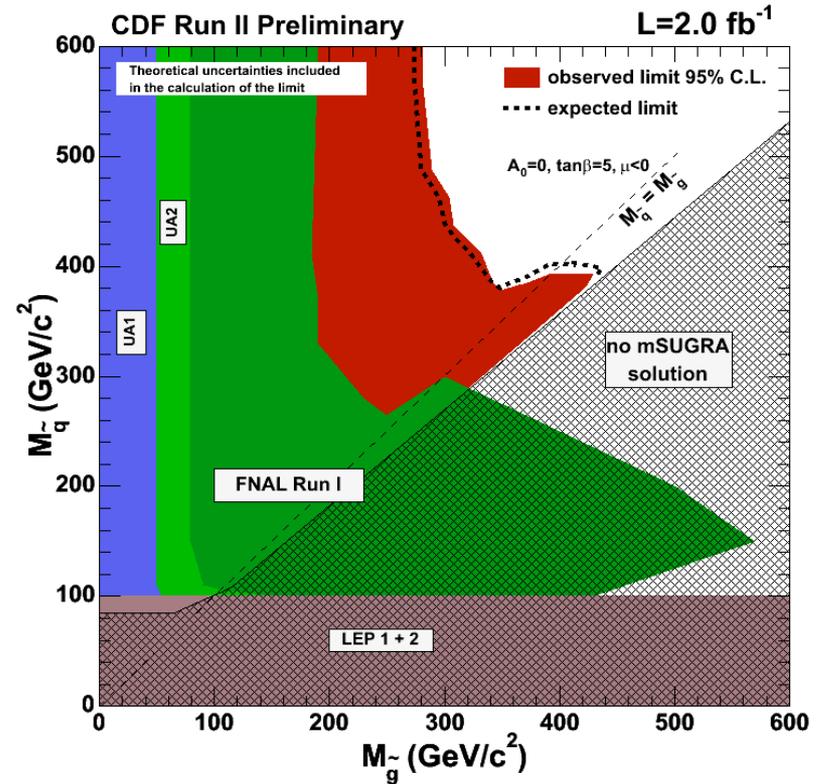
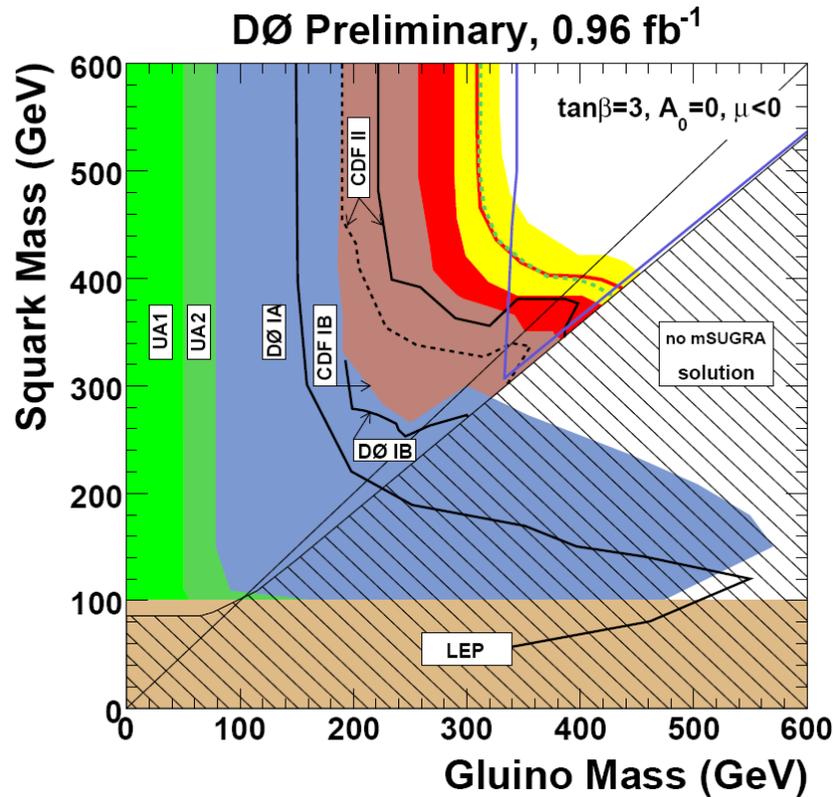
Observed events in data:

Region	Observed data
4-jets	45
3-jets	38
2-jets	18

No excess above background from Standard Model processes

→ No evidence for SUSY (yet) → Set limits on masses of SUSY particles

Excluded regions in the m(squark) vs. m(gluino) plane



Exclusion limits

(incl. systematic uncertainties)*:

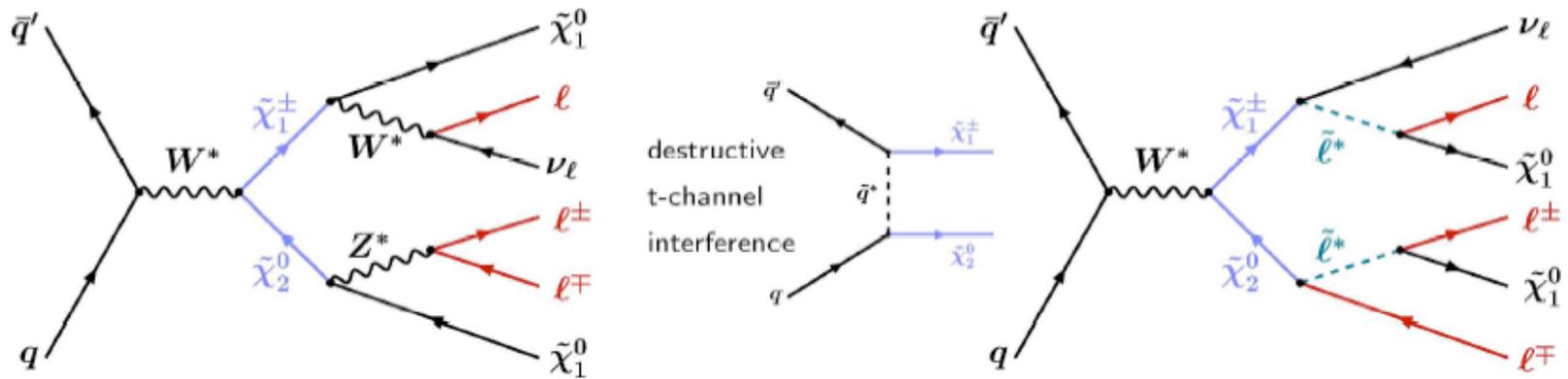
$$m(\text{gluino}) > 290 \text{ GeV}/c^2$$

$$m(\text{squark}) > 375 \text{ GeV}/c^2$$

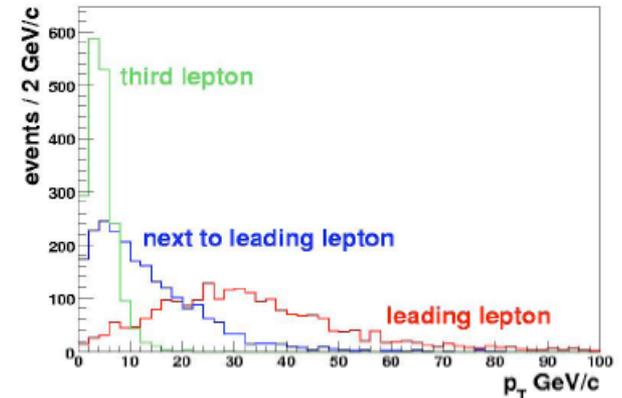
)* uncertainties from structure functions, change of renormalization and factorization scale μ by a factor of 2, NLO calculation, default choice: $\mu = m(\text{gluino}), m(\text{squark})$ or $\frac{1}{2}(m(\text{gluino})+m(\text{squark}))$ for gg, qq, qg production

Search for Charginos and Neutralinos - the tri-lepton channel-

- Gaugino pair production via electroweak processes
(small cross sections, $\sim 0.1 - 0.5$ pb, however, small expected background)



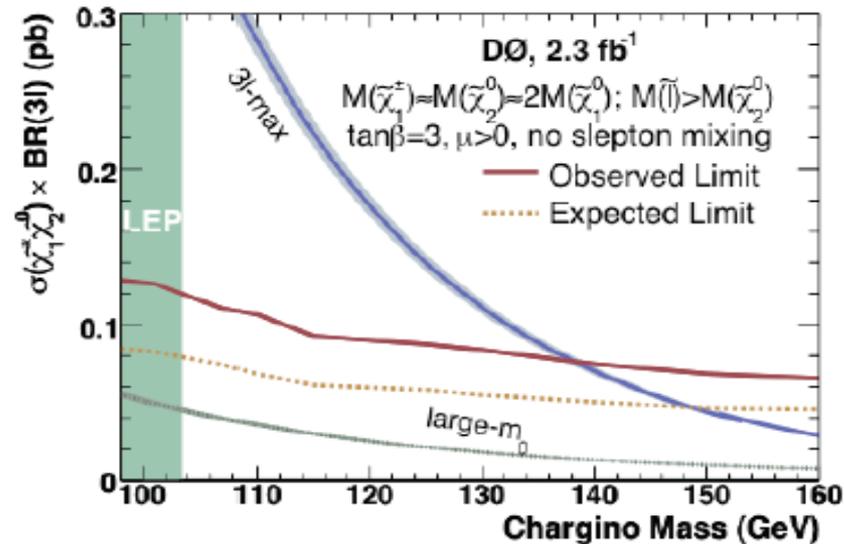
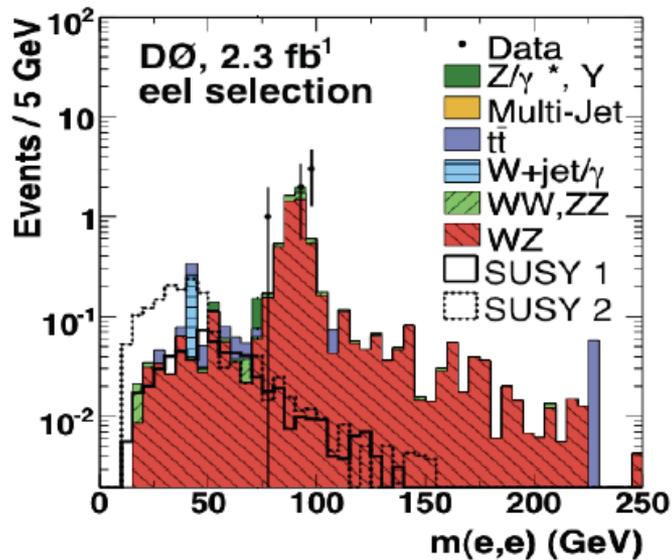
- For small gaugino masses (~ 100 GeV/ c^2) one needs to be sensitive to low P_T leptons



Analysis:

- Search for different ($\ell\ell\ell$) + like-sign $\mu\mu$ final states with missing transverse momentum
- In order to gain efficiency, no lepton identification is required for the 3rd lepton, select: two identified leptons + a track with $p_T > 4$ GeV/c

mSUGRA interpretation



For specific scenarios: sensitivity / limits above LEP limits;
 e.g., $M(\chi^\pm) > 140$ GeV/c² for the 3l-max scenario